

Afterglow Light
Pattern
380,000 yrs.

Dark Ages

Development of
Galaxies, Planets, etc.

Dark Energy
Accelerated Expansion

Beyond Standard Model Big-Bang Nucleosynthesis: Fundamental Constants

Ermal Rrapaj, Jianping Lai, Alexander Bartl

Quantum
Fluctuations

1st Stars
about 400 million yrs.

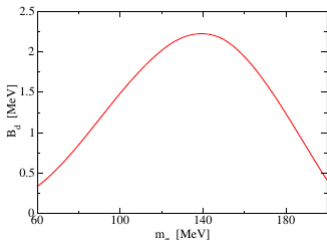
Big Bang Expansion

13.7 billion years

- BBN important for synthesis of light elements and evolution of universe
- Element abundances as probes of fundamental physics
- Impact of variations of Λ_{QCD} , G , μ_n on BBN

- 1 Background information
 - Λ_{QCD} and m_n , deuteron binding energy, neutron lifetime
 - $m + n \rightarrow d + \gamma$ reaction rate from μ_n
 - Graviational Constant
- 2 Calculations
- 3 Results
 - Abundance Plots
- 4 Conclusions

Impact of Λ_{QCD}

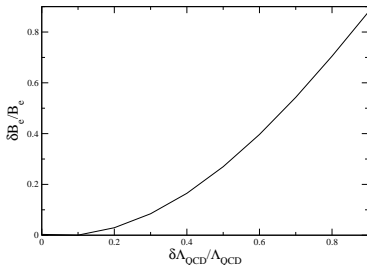
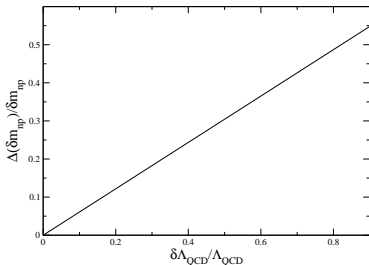


B_D vs. m_π at NLO χEFT , Savage et al, 2002

Gail McLaughlin et al, 2003

- $\Delta(\delta m_{np}) = -\left(\frac{\Delta\alpha}{\alpha} + \frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}\right)\alpha M_{elm}$
- $\frac{\Delta\alpha}{\alpha} \approx 1/30 \frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}$
- $\alpha M_{elm} \approx 0.76 \text{ MeV}$
- $\frac{\Delta m_\pi}{m_\pi} = \frac{1}{2} \left(\frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}}\right)$

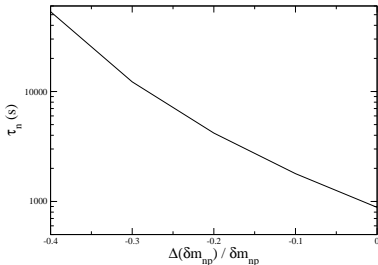
(Absolute values of differences on y-axis)



Impact Of Λ_{QCD} , G

Dependence of neutron lifetime:

$$\tau_n = \frac{(G_F \cos \theta_c)^2}{2\pi^3} m_e^5 (1 + g_A^2) F(\Delta m_{np})$$
$$F(x) = \frac{1}{15} (2x^4 - 9x^2 - 8) \sqrt{x^2 - 1} + x \log(x + \sqrt{x^2 - 1})$$



Friedman Equations:

$$H^2 + \frac{k}{a^2} = \frac{8\pi G}{3} \rho + \frac{\Lambda_{vac}}{3}$$
$$H^2 + \frac{dH}{dt} = -\frac{4\pi G}{3} (\rho + 3p) + \frac{\Lambda_{vac}}{3}$$

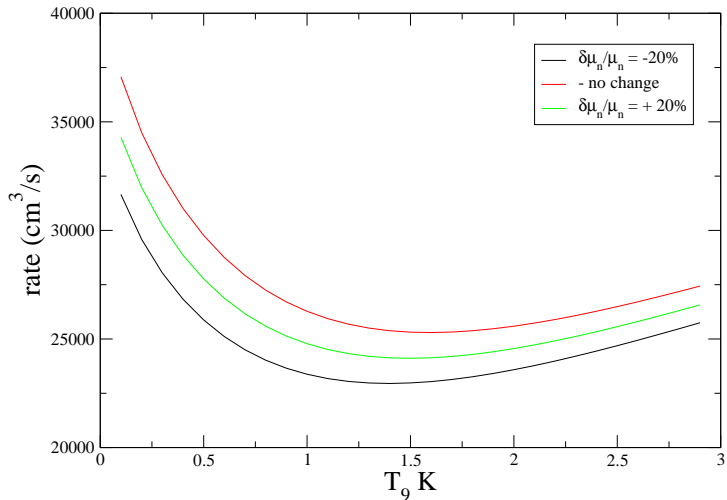
Radiation Dominated, $k \approx 0$, $\Lambda_{vac} \approx 0$:

$$\rho = 3p; \quad \rho = \frac{3}{32} \pi G t^2$$
$$\implies H = \frac{\pi}{2} G t$$

Impact of μ_n on neutron capture rate on protons

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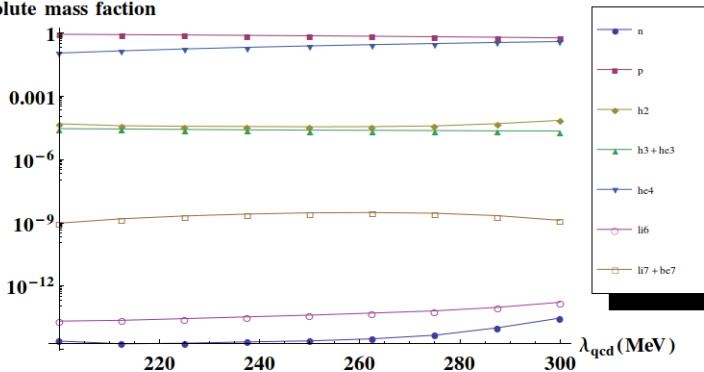


Results : Mass fraction as function of Λ_{QCD} , part 1

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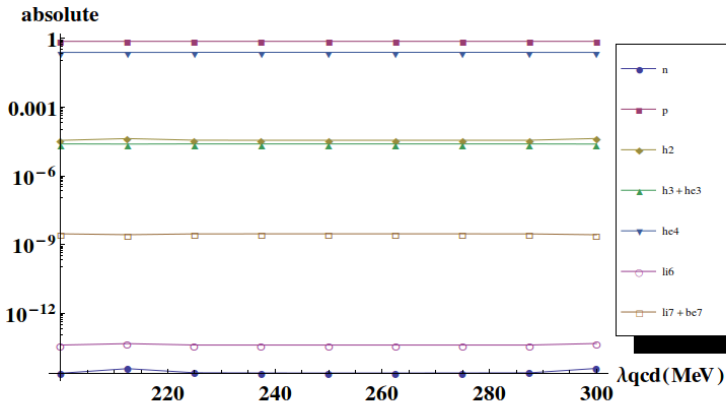
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absolute mass fraction



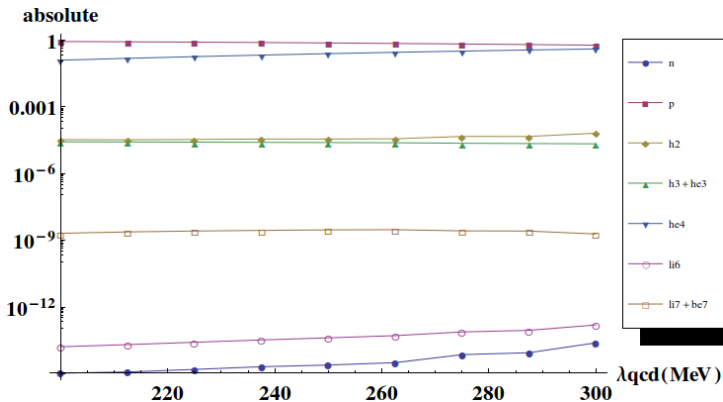
Results : Mass fraction as function of Λ_{QCD} , part 2

Neutron Mass Change Only



Results : Mass fraction as function of Λ_{QCD} , part 3

τ_n Change only

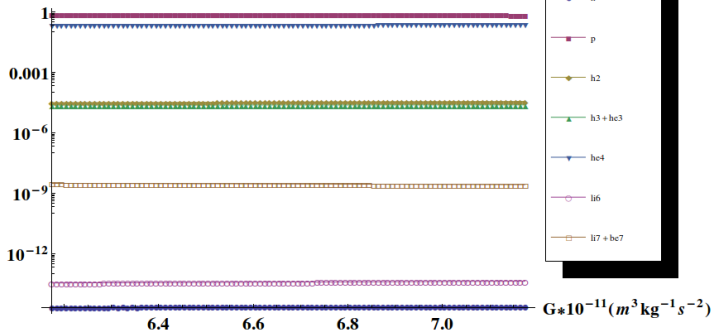


Results : Mass fraction as function of G , part 1

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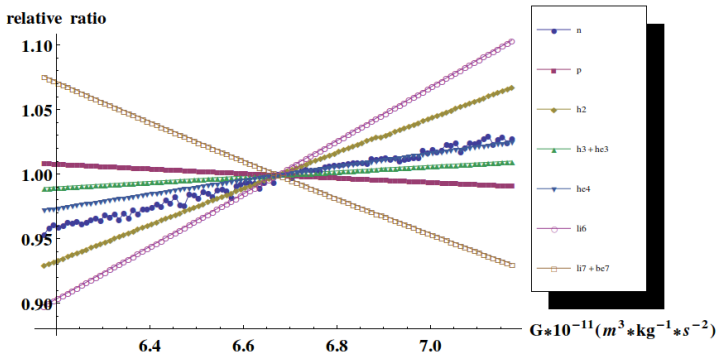
absolute ratio



Results : Mass fraction as function of G , part 2

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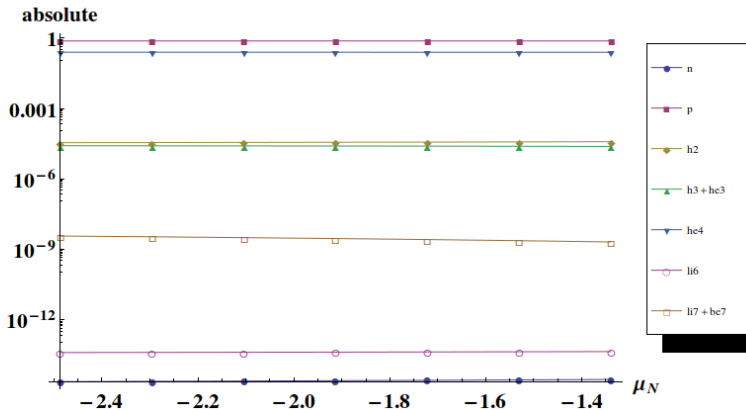
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Results : Mass fraction as function of μ_n , part 1

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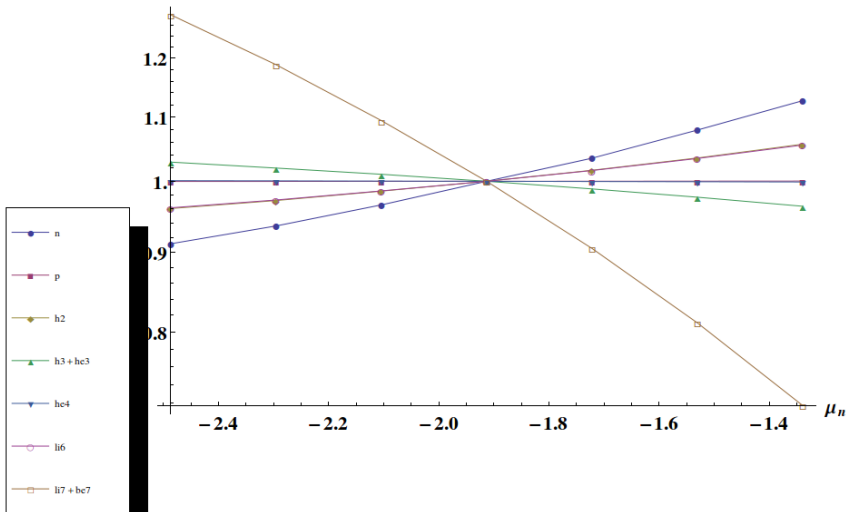


Results : Mass fraction as function of μ_n , part 2

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relative mass fraction



- We investigated the dependence of BBN abundances on the values of fundamental constants.
 - Λ_{QCD} has a significant impact especially on the Helium-4 abundance, mainly through the change to the neutron decay time.
 - The lithium abundance is highly sensitive to the value of the gravitational constant G .
 - Changing μ_n decreases the rate of production of deuteron
-
- A change in Λ_{QCD} would probably affect more than the $p(n, \gamma)d$ rate and free neutron decay. Yet, modifying more interactions is beyond the scope of this 4-day project.