

# I) Motivation for Microscopically-based EDFs

## a) Successes of EDF (Skyrme etc.)

- \* Computational simplicity (mass tables in a day, global studies a-la PH Heenen's talks  
fusion barriers/energy surfaces, etc.)
- \* rms error  $\sim 1$  MeV (versus  $\frac{1 \text{ MeV}}{A}$  for ab-initio in limited medium mass)
- \* Access open shell/deformed nuclei
- \* Show figures!

## b) Failures/Shortcomings $\Rightarrow$ (why insights from ab-initio might help)

- \* stuck @  $\sim 1$  MeV brick wall despite influx of sophisticated optimization methods
- \* performance for spectroscopy not as solid
- \* poorly constrained isovector +  $\lambda$ -odd couplings
- \* Loss of predictive power away from mass region  
/observables fitted to (show, e.g.,  $S_{n \text{ figs in } S_n}$ )

## \*\* $E[\rho]$ too simple!

- \* Why only bilinear in  $\rho$ 's  
w/ the lone  $\rho^{2+\lambda}$  term?

e.g., 3NF play big role in ab-initio ( $\Rightarrow$  tri-linear, at least)

e.g., Even toy model systems (e.g., dilute gas of fermions in a trap  
where one can constructively interesting w/ short-range forces)  
build  $E_{\text{HK}}[\rho]$  has multiple non-analytic (non-integers)  
terms of the form  $\rho^\lambda$

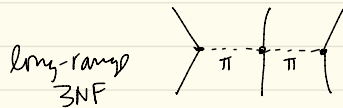
- \* In DME, even @ HF level see non-analytic behavior  
due to finite range

# E[P] too simple (cont'd)

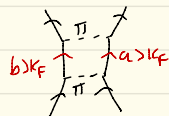
\* ab-initio (Show SM + GFMC plots) calculations  $\Rightarrow$  2 Very different sources of LS physics



Short-range  
NN LS-force



long-range  
3NF



iterated  
 $\pi$ -exchange



Consistent w/ Skyrme

$$iW_0 (\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{k}' \times \delta(\vec{r}_1 - \vec{r}_2) \vec{k}$$



How can zero range  
Skyrme capture this?

\* Pion-exchanges ( $\gamma$ ,  $\rho$ ,  $\omega$ ,  $\dots$ ) have very rich spin/isospin/space dependence. Where are they in Skyrme? Could putting them in solve/ameliorate the poorly constrained isovector &  $t$ -odd couplings?

\* Aesthetics: Can we at least have indirect link to QCD?  $\chi$ -EFT theory of forces has this link. [**\* show EFT figure**]

$\Rightarrow$  Can we connect E[P] to  $\chi$ -EFT 2- + 3N interactions?

# Thoughts on Coulomb DFT (see Perdew reference posted)

LDA: Do ab-initio calc. of the electron gas over wide range of  $\rho$

\* extract  $\Rightarrow \frac{E_{xc}^{gas}}{A} = \epsilon_{xc}(\rho)$  parameterizing dep. on  $\rho$ .

\* treat finite ITH system like it's a chunk of  $e^-$  gas locally (LDA)

$$\Rightarrow E_{xc}^{LDA} = \int d^3r \rho(\vec{r}) \epsilon_{xc}(\rho(\vec{r}))$$

$$V_{xc}^{LDA}(\vec{r}) = \frac{\delta E_{xc}^{LDA}}{\delta \rho(\vec{r})} = \epsilon_{xc}(\rho(\vec{r})) + \rho(\vec{r}) \frac{\partial \epsilon_{xc}(\rho(\vec{r}))}{\partial \rho}$$

## \* Gradient corrections

\* early attempts using linear response failed (corrections made things worse)

$\Downarrow$

$$E_{\nabla}[P] \sim \int d^3r C |\nabla P(\vec{r})|^2$$

huge correction.

## \* Generalized Gradient Approx (GGA)

- empirical form w/eg., renormalized gradients in denominator

e.g.,  $\int d^3r \rho(\vec{r}) \epsilon_c(\rho) \left[ 1 - \frac{\beta |\nabla P|^2}{\nu \rho \epsilon_c(\rho)} \right]^{-\nu}$  (M+Becke)

$\nu$  = fit parameter

\* later versions by Perdew et al. choose complicated form, impose sum rule + general scaling relations

## Perdew et al.: Non-empirical GGA functionals

- \* start w/ some ansatz for re-summed gradients
- \* impose exact relations (sum rules, scaling properties, etc.)  
to essentially eliminate any fit parameters in the ansatz

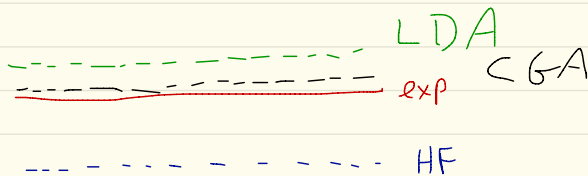
Comment: To my knowledge, these constraints using exact relations follow from the simplicity of  $V_{\text{Coulomb}}$ .

Probably can't do the same in nuclear physics  
do to extreme complexity of  $V_{NN} + V_{3N}$ .

However, see e.g. the work of Dobaczewski  
using gauge-invariance etc. to find relations  
between couplings.

Question: Does Coulomb DFT work so well in part  
because XC effects small?

e.g., HF is good starting point



With soft EFT + RG-evolved  $2N+3N$  interactions,  
HF becomes a not crazy (still not quantitative)  
starting point, w/ bulk of correlation energy coming  
in low order MBPT/BHP type correlations

## II) Skyrme from BHF

\* BHF crash review

{ upshot, signs look like HF

$$E^{\text{BHF}} = \sum_{i=1}^A \langle i | t | i \rangle + \frac{1}{2} \sum_{ij=1}^A \langle ij | G(\omega = \epsilon_i + \epsilon_j) | ij \rangle$$

$$G(\omega) = V + V \frac{Q}{\omega - t} G \quad Q = \sum_{ab} |ab\rangle \langle ab| + \sum_{ai} |ai\rangle \langle ai|$$

$$\Sigma_i = t_{ii} + \sum_j G_{ijij}(\omega = \epsilon_i + \epsilon_j)$$

\* for practice to make BHF as much as possible,

$$(1) G(\omega = \epsilon_i + \epsilon_j) \Rightarrow G(\epsilon_{av}) \quad \sum_{ij} \langle ij | G(\epsilon_{av}) | ij \rangle \text{ in } P\text{-matrix}$$

$$(2) \langle \vec{r}_1 \vec{r}_2 | G | \vec{r}_3 \vec{r}_4 \rangle = \delta(\vec{r} - \vec{r}') G_{ijij}^{\vec{r}}(\vec{r}, \vec{r}') \approx \delta(\vec{R} - \vec{R}') \delta(\vec{r} - \vec{r}') G_{ijij}^{\vec{R}}$$

$$\Rightarrow E_{int}^{\text{BHF}} \sim \int d\vec{r} d\vec{r}' \rho^2(\vec{r} \pm \frac{1}{2}\vec{\epsilon}, \vec{r} - \frac{1}{2}\vec{\epsilon}) G(\vec{r})$$

\* Skyrme  $\neq$  zero range exp. of  $G(r)$

\* What sets fall off of  $\rho(\vec{r} \pm \frac{1}{2}\vec{\epsilon}, \vec{r} - \frac{1}{2}\vec{\epsilon})$ ? } nm example  
140 example  $\Rightarrow$  if  $\rho(\vec{r})$  does.

# Crash Course on BHF

HF:  $\Psi(1, \dots, A) \approx \mathcal{A}(\phi_i^{(1)} \phi_j^{(2)} \dots)$  HF orbitals



Naive HF won't even bind  
due to S.R. repulsion

- however, as Peter said,  
nuclei sit here

- likewise  $\lambda_{\text{MFP}} \sim \text{Size of nucleus}$

Nucleus is  
"dilute"

$\Rightarrow$  suggests the next level of approx, BHF or Indep. Pair Approx.

$$\Psi(1, 2, \dots, A) \approx N \sum_{k\ell} \mathcal{A}(\underbrace{\Psi_{k\ell}(1, 2)}_{\text{correlated pair}} \phi_m^{(3)} \dots \phi_n^{(A)}) \underbrace{(T(i) + U_{\text{MF}}(i))}_{H_0(i)} \phi_i^{(j)} = \sum_{\mathbf{q}} \epsilon_{\mathbf{q}} \phi_{\mathbf{q}}^{(i)}$$

Correlated pair:  $(H_0(1) + H_0(2) + V_{12}) \Psi_{k\ell}(1, 2) = e_{k\ell} \Psi_{k\ell}(1, 2)$

expand:  $|\Psi_{k\ell}\rangle = |\phi_k \phi_\ell\rangle + \sum_{a, b \in \mathcal{F}} |\phi_a \phi_b\rangle \langle \phi_a \phi_b | \Psi_{k\ell}\rangle$

(add  $|\phi_a \phi_b\rangle, |\phi_b \phi_a\rangle$  terms infinite  
nuclei, but I ignore them)

$$|\Psi_{k\ell}\rangle = |\phi_k \phi_\ell\rangle + \frac{Q}{e_{k\ell} - H_0} V |\Psi_{k\ell}\rangle$$

$$Q = \sum_{a, b \in \mathcal{F}} |\langle \phi_a \phi_b | \phi_k \phi_\ell \rangle|$$

Bethe-Goldstone eqn (Pauli-Blocked Lippman-Schwinger)

G-matrix: Just like  $\hat{V}|\Psi\rangle = \hat{T}|\Phi\rangle$  in LS eqn in scattering theory,

$$\underline{Def} \quad \hat{V}|\Psi_{k\ell}\rangle = G(\omega = \epsilon_{k\ell})|\Phi_k\Phi_\ell\rangle$$

$\Rightarrow$  Bethe Goldstone eqn then gives

$$G(\omega) = \mathcal{V} + \mathcal{V} \frac{Q}{\omega - H_0} G(\omega)$$

Can show w/ these definitions

$$E^{BHF} = \sum_{\lambda=1}^A \langle \phi_\lambda | T | \phi_\lambda \rangle + \frac{1}{2} \sum_{\lambda, j=1}^A \langle \phi_\lambda \phi_j | G(\epsilon_\lambda + \epsilon_j) | \phi_\lambda \phi_j \rangle$$

$$\Sigma_\lambda = \langle \phi_\lambda | T | \phi_\lambda \rangle + \sum_{j=1}^A \langle \phi_\lambda \phi_j | G(\epsilon_\lambda + \epsilon_j) | \phi_\lambda \phi_j \rangle$$

Comments: ① looks like HF  $V \rightarrow G(\epsilon_i + \epsilon_j)$  [Doubly self-consistent though]

$$\left. \begin{aligned} \text{② Even if } \langle \vec{r}_1 \vec{r}_2 | V | \vec{r}_3 \vec{r}_4 \rangle &= \delta(\vec{r}_1 - \vec{r}_3) \delta(\vec{r}_2 - \vec{r}_4) V(\vec{r}) \\ &= \delta(\vec{R}_{12} - \vec{R}_{34}) \delta(\vec{r}_{12} - \vec{r}_{34}) V(\vec{r}_{12}) \end{aligned} \right\} \text{Local}$$

$$\langle \vec{r}_1 \vec{r}_2 | G | \vec{r}_3 \vec{r}_4 \rangle = \delta(\vec{R}_{12} - \vec{R}_{34}) \underbrace{G(\vec{r}_{12}, \vec{r}_{34})}_{\substack{\text{Non-local,} \\ \text{parametric (needs)} \\ \text{com dependence}}}$$

③ Energy dependence prevents  
from writing  $E_{BHF}$  in terms  
of  $\rho(\vec{r}_1, \vec{r}_2), \rho(\vec{r})$ .

## Comments (Cont'd):

④ BHF replaced by more sophisticated theories in practice (CC, IMSRG, ...)  
but all such methods dominated by BHF-type correlations

⑤ Notice no prescription given for how to compute  $\langle \epsilon_q \rangle_{E_F}$

Technical point: different prescriptions exist. Ⓞ Using old-fashioned "hard"  $V$  results can be very sensitive. Claim this goes away w/ softer interactions now in vogue.

## Making BHF look more like HF

1) avg out  $E$ -dependence, e.g.  $G(\epsilon_i + \epsilon_j) \rightarrow G(\bar{\epsilon}_{\text{pair}})$

2) avg out non-locality  $\Rightarrow G(\vec{r}_{12}, \vec{r}_{34}, \bar{\epsilon}) \rightarrow f(\vec{r}_{12} - \vec{r}_{34})$   
& com dep  $G(\vec{r}_{12})$



Then,

$$E_{\text{BHF}} \sim \frac{1}{2} \sum_{ij} \langle ij | f(r) | ij \rangle$$
$$\sim \frac{1}{2} \iint d\vec{r} d\vec{r}' \rho(\vec{r} + \frac{\vec{r}}{2}) \rho(\vec{r} - \frac{\vec{r}}{2}) G(r)$$
$$- \frac{1}{2} \iint d\vec{R} d\vec{r} \rho^2(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) G(r)$$

[reminds,  $G(r)$  is  $\rho$ -dependent!]



Final point Re: BHF

Correlations are "short-range"

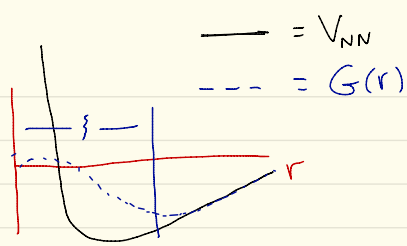
$\xi$  = "healing distance"

$$\sim \frac{\pi}{k_F}$$

crudely speaking;

$$G(r) \sim \Theta(r-\xi) V(r) + \Theta(\xi-r) [c_0 \delta(r) + c_2 \nabla^2 \delta(r) + \dots]$$

Will make use of this later on...



\*  $\hat{V}_{\text{Skyrme}}$  is NOT a zero-range expansion of  $V_{NN}(r)$  (or  $G_{NN}(r)$ )

\* It is sometimes incorrectly said that  $\hat{V}_{\text{Sk}}$  is a zero-range approx. to  $V_{NN}/G_{NN}$ , which are said to be "short ranged"

- Range of NN-force controlled by  $m_{\pi}$ ; i.e.,  $V(r) \sim \frac{g e^{-m_{\pi} r}}{r}$

- Under what conditions can we take  $V(r) \sim C_0 \delta(\vec{r}) + C_2 \nabla^2 \delta(\vec{r}) + \dots$  like  $\hat{V}_{\text{Sk}}$ ?

- Go over to momentum space:  $\tilde{V}(\vec{q}) = \int \frac{d^3 r}{(2\pi)^3} e^{i\vec{q}\cdot\vec{r}} V(r)$

$$\propto \frac{g}{\vec{q}^2 + m_{\pi}^2}$$

\* at  $q \lesssim m_{\pi}$ :  $\tilde{V}(\vec{q}) \approx C_0 + C_2 q^2 + \dots$

$\Downarrow$  FT back to  $r$ -space

$$V(r) \approx C_0 \delta^{(3)}(\vec{r}) + C_2 \nabla^2 \delta^{(3)}(\vec{r}) + \dots$$

$\Rightarrow$  Can approximate  $V(r) \sim$  zero-range only for cases where you probe momenta  $q \lesssim m_{\pi}$

But

$$\left. \begin{aligned} E_{\text{ind}}^{\text{HF}} &\sim \sum_{ij=1}^A \langle ij | V | ij \rangle \\ E_{\text{ind}}^{\text{BHF}} &\sim \sum_{ij=1}^A \langle ij | G | ij \rangle \end{aligned} \right\} \begin{array}{l} \text{probe momenta } \mathcal{O}(k_F) \\ k_F \sim 1.3 \text{ fm}^{-1} \\ m_{\pi} \sim 0.7 \text{ fm}^{-1} \end{array} \quad \Bigg\|$$

$\Rightarrow$  Misleading to say  $V_{SK} \sim$  zero range expansion of  $V_{NN}(r)/f_{NN}(r)$

Remark: Approximating  $V(\vec{r}) \sim C_0 \delta(\vec{r}) + C_2 \nabla^2 \delta(\vec{r}) + \dots$  equivalent to doing Taylor expansion about the diagonal of the density matrix

ex:  $E_x = \frac{1}{2} \int d\vec{r}_1 d\vec{r}_2 \rho^2(\vec{r}_1, \vec{r}_2) V(\vec{r}_1 - \vec{r}_2) = \frac{1}{2} \int d\vec{R} d\vec{r} \rho^2(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) V(r)$

let  $\rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \sim \rho(\vec{R}, \vec{R}) + \underbrace{\vec{r}_i \cdot \nabla_i \rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2})}_{\substack{\text{0 in even-even} \\ \text{by } T\text{-reversal}}} \Big|_{\vec{r}=0} + \frac{1}{2} \vec{r}_i \nabla_i \vec{r}_j \nabla_i \nabla_j \rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \Big|_{\vec{r}=0} + \dots$

eg.: LO term  $\frac{1}{2} \int d\vec{R} d\vec{r} \rho^2(\vec{R}) V(r) = \frac{1}{2} \int d\vec{R} \rho^2(\vec{R}) \int d\vec{r} V(r)$

from expanding DM

$\vec{q}=0$  term of  $\tilde{V}(\vec{q})$   
(i.e., 1st term in Taylor expanding  $\tilde{V}(\vec{q}) \approx \tilde{V}(0) + \dots$ )

\* You can verify the  $\vec{r} \cdot \vec{\nabla} \rho$  term vanishes

in even-even system due to time reversal, while the

cross-terms  $\rho(\vec{R}) \frac{1}{2} \vec{r}_i \cdot \vec{\nabla} \vec{r}_j \nabla_i \nabla_j \rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \Big|_{\vec{r}=0}$  give the same

term as for the 2<sup>nd</sup> term of expanding  $\tilde{V}(q) \approx \tilde{V}(0) + q^2 \tilde{V}'(q^2=0)$

Question: Might it be possible to find a resummed Taylor expansion that "factorizes" the non-locality (r-dependence) of density matrix?

$$\rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \approx \sum_n \prod_{\vec{r}_n} V^{\vec{r}_n} \rho(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \Big|_{\vec{r}=0}$$

all we require at this point is that  $\prod_{2n} V^{\vec{r}_n}$  has correct small-r behavior to match Taylor exp.

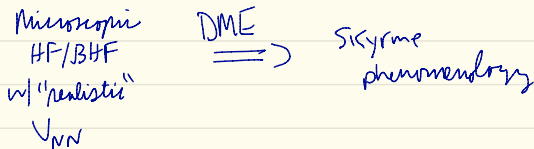
$\Rightarrow$  Non-local Fock energy calculated w/ a realistic NN pot. then has a form that resembles what  $V_{\text{Skyrme}}$  gives

$$E_x = \frac{1}{2} \left( \int d\vec{r} d\vec{r}' V(r) \rho^2(\vec{r} + \frac{\vec{r}}{2}, \vec{r} - \frac{\vec{r}}{2}) \right) \approx \frac{1}{2} \sum_{n,m} \int d\vec{R} \left( \underbrace{\nabla^{2n} \rho(\vec{R})}_{\text{local densities + their gradients}} \right) \nabla^{2m} \rho(\vec{R}) \times \int d\vec{r} \underbrace{\rho_{2n}(r) \rho_{2m}(r) V(r)}_{\text{coupling from underlying NN}}$$

$$= \int d\vec{R} \left[ \rho^2(\vec{R}) g_{00} + \rho(\vec{R}) \nabla^2 \rho(\vec{R}) g_{02} + \dots \right]$$

\* OK, so we see that "factorizing" the non-locality (i.e.  $r$ -dependence) of  $\rho(\vec{r} + \frac{\vec{r}}{2}, \vec{r} - \frac{\vec{r}}{2})$  allows us to map non-local Fock-energy into a local Skyrme-like EDF!

\* We call this expansion the "Density matrix expansion" (DME). I haven't told you how to do this yet (Moreover, there is no 1 unique prescription), but this gives us a hint that the Skyrme EDF can be understood as performing the DME on a microscopic HF/BHF calculation



\* Before we go over a couple recipes for the DME, need to understand what sets the scale for the falloff in  $r$ -direction of  $\rho(\vec{r} + \frac{\vec{r}}{2}, \vec{r} - \frac{\vec{r}}{2})$

ex1: infinite homogeneous matter; HF sp wf's = plane waves  $\phi_{\vec{k}}(\vec{r}) = \frac{e^{i\vec{k}\cdot\vec{r}}}{\sqrt{V}} \chi_{\sigma}$

$$\rho(\vec{r}_1, \vec{r}_2) = 2 \sum_{\vec{k} \in \text{KF}} \phi_{\vec{k}}^*(\vec{r}_2) \phi_{\vec{k}}(\vec{r}_1) = \frac{2}{V} \sum_{\vec{k} \in \text{KF}} e^{i\vec{k}\cdot(\vec{r}_1 - \vec{r}_2)} \xrightarrow{V \rightarrow \infty} 2 \int_{\frac{\text{KF}}{(2\pi)^3} d^3k} e^{i\vec{k}\cdot\vec{r}}$$

$$\Rightarrow \rho_{NM}(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) = \frac{2 \cdot 4\pi}{(2\pi)^3} \int_0^{k_F} k^2 dk j_0(kr) = \frac{1}{\pi^2} k_F^3 \int_0^1 \bar{k}^2 d\bar{k} j_0(\bar{k}r) \quad \begin{matrix} \bar{k} = \frac{k}{k_F} \\ \bar{r} = r k_F \end{matrix}$$

$$\frac{1}{3\pi^2} k_F^3 = \rho \Rightarrow$$

$$= 3\rho \int_0^1 \bar{k}^2 d\bar{k} j_0(\bar{k}r)$$

$$= 3\rho \frac{j_1(k_F r)}{k_F r}$$

$$\rho_{NM}(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) = \rho \rho_{SL}(k_F r)$$

$$\rho_{SL}(k_F r) = \frac{3 j_1(k_F r)}{k_F r}$$

i.e.,  $k_F$  controls falloff in  $r$ .

Comment: The Slater Approximation mentioned by Nicolas for treating Coulomb exchange in Skyrme calculations derived by taking  $\rho_{NM}(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \approx \rho(\vec{R}) \rho_{SL}(k_F(\vec{R})r)$  ("local density approx" or LDA)

# Phase-Space Averaging (PSA) Formulation of DME (Geckeler, SICB/Duguet)

\* Start w/ spin/isospin decomposition of density matrix

$$\tilde{\rho}(\vec{r}_1, \vec{r}_2) = \frac{1}{4} \sum_m \sum_\nu \rho_{m\nu}(\vec{r}_1, \vec{r}_2) \hat{\sigma}_m \hat{\tau}_\nu \quad \hat{\sigma}_m = (\mathbb{1}, \vec{\sigma})$$

Matrix in  $\sigma\tau$  space

$$\rho_{m\nu}(\vec{r}_1, \vec{r}_2) = \text{Tr}_\sigma \text{Tr}_\tau \tilde{\rho}(\vec{r}_1, \vec{r}_2) \hat{\sigma}_m \hat{\tau}_\nu$$

\* Formally expand out the non-locality of  $\rho_{m\nu}(\vec{r}_1, \vec{r}_2) = \rho_{m\nu}(\vec{R} + \frac{\vec{r}_1 - \vec{r}_2}{2}, \vec{R} - \frac{\vec{r}_1 - \vec{r}_2}{2})$

$$\rho_{m\nu}(\vec{R} + \frac{\vec{r}_1 - \vec{r}_2}{2}, \vec{R} - \frac{\vec{r}_1 - \vec{r}_2}{2}) = e^{\frac{\vec{r}_1 - \vec{r}_2}{2} \cdot \vec{\nabla}_r} \rho_{m\nu}(\vec{R} + \frac{\vec{r}_1 - \vec{r}_2}{2}, \vec{R} - \frac{\vec{r}_1 - \vec{r}_2}{2}) \Big|_{\vec{r}_1 = \vec{r}_2 = \vec{R}} = e^{\frac{\vec{r}_1 - \vec{r}_2}{2} \cdot \vec{\nabla}_r} \rho_{m\nu}(\vec{r}_1, \vec{r}_2) \Big|_{\vec{r}_1 = \vec{r}_2 = \vec{R}}$$

\* Multiply by  $1 = e^{i\vec{r} \cdot \vec{K}} e^{-i\vec{r} \cdot \vec{K}} \Rightarrow$

$$\rho_{m\nu}(\vec{R} + \frac{\vec{r}_1 - \vec{r}_2}{2}, \vec{R} - \frac{\vec{r}_1 - \vec{r}_2}{2}) = e^{i\vec{r} \cdot \vec{K}} e^{\frac{\vec{r}_1 - \vec{r}_2}{2} \cdot \vec{\nabla}_r} \rho_{m\nu}(\vec{r}_1, \vec{r}_2) \Big|_{\vec{r}_1 = \vec{r}_2 = \vec{R}}$$

$$\approx e^{i\vec{r} \cdot \vec{K}} \left[ 1 + \vec{r} \cdot \left( \frac{\vec{\nabla}_2}{2} - i\vec{K} \right) + \frac{1}{2} \left( \vec{r} \cdot \left( \frac{\vec{\nabla}_2}{2} - i\vec{K} \right) \right)^2 \right] \rho_{m\nu}(\vec{r}_1, \vec{r}_2) \Big|_{\vec{R}} \quad \otimes$$

\* physically,  $\vec{K} \sim$  some avg. rel momentum in nucleus

\* Assume we have some model local momentum distribution  $g(\vec{R}, \vec{K})$

$\Rightarrow$  Want to average  $\otimes$  over  $g(\vec{R}, \vec{K})$

$$D_L: \Pi_n(\vec{r}, \vec{R}) = \frac{\int d\vec{R} e^{i\vec{r}\cdot\vec{R}} (\vec{r}\cdot\vec{R})^n g(\vec{R}, \vec{R})}{\int d\vec{R} g(\vec{R}, \vec{R})}$$

$$\vec{J}_{m\nu}(\vec{R}) = -\frac{i}{2} (\vec{\nabla}_1 - \vec{\nabla}_2) \mathcal{P}_{m\nu}(\vec{r}_1, \vec{r}_2) \Big|_{\vec{r}_1 = \vec{r}_2 = \vec{R}}$$

$$\mathcal{T}_{ab, m\nu}(\vec{R}) = \vec{\nabla}_1^a \vec{\nabla}_2^b \mathcal{P}_{m\nu}(\vec{r}_1, \vec{r}_2) \Big|_{\vec{r}_1 = \vec{r}_2 = \vec{R}}$$

Spatial indices (KE Tension density)

$$\Rightarrow \mathcal{P}_{m\nu}(\vec{r}_1, \vec{r}_2) \approx \left[ \Pi_0 - i\Pi_1 - \frac{\Pi_2}{2} \right] \mathcal{P}_{m\nu}(\vec{R}) + i \left[ \Pi_0 - i\Pi_1 \right] \sum_{a=1}^3 r_a \dot{\mathcal{J}}_{a, m\nu}(\vec{R}) + \frac{\Pi_0}{2} \sum_{a,b=1}^3 r_a r_b \left[ \frac{1}{4} \nabla_a \nabla_b \mathcal{P}_{m\nu}(\vec{R}) - \mathcal{T}_{ab, m\nu}(\vec{R}) \right]$$

⊗

\* Eg:  $g(\vec{R}, \vec{R}) = \Theta(k_{FR} - |\vec{R}|)$  (Phase space of  $\infty$ -matter)

$$\Rightarrow \Pi_0(k_{FR}) = \frac{3\dot{\mathcal{J}}_1(k_{FR})}{k_{FR}} = \mathcal{P}_{SL}(k_{FR}) \approx 1 + \mathcal{O}(k_{FR})^2$$

$$\Pi_1(k_{FR}) = -i3\dot{\mathcal{J}}_0(k_{FR}) + i9 \frac{\dot{\mathcal{J}}_1(k_{FR})}{k_{FR}} \approx i \frac{(k_{FR})^2}{5} + \mathcal{O}(k_{FR})^4$$

$$\Pi_2(k_{FR}) = 15\dot{\mathcal{J}}_0 - 36 \frac{\dot{\mathcal{J}}_1}{k_{FR}} - 3\omega k_{FR} \approx \frac{(k_{FR})^2}{5} + \mathcal{O}(k_{FR})^4$$

\* We further re-order ⊗ by counting  $k_F \sim \nabla$

$$\rho_{mv}(\vec{r}_1, \vec{r}_2) \approx \Pi_0(k_F r) \rho_{mv}(\vec{R}) + i \Pi_1(k_F r) \sum_{a=1}^3 r_a j_{a,mv}(\vec{R}) + \frac{\Pi_2(k_F r)}{2} \sum_{a,b} r_a r_b \left[ \frac{1}{4} \nabla_a \nabla_b \rho_{mv}(\vec{R}) - \zeta_{ab,mv}(\vec{R}) + d_{ab} \frac{k_F^2}{5} \rho_{mv}(\vec{R}) \right]$$

$$\equiv \Pi_0(k_F r) \rho_{mv}(\vec{R}) + i \Pi_1(k_F r) \sum_{a=1}^3 r_a j_{a,mv}(\vec{R}) + \frac{\Pi_2}{2} \sum_{a,b} r_a r_b \left[ \frac{1}{4} \nabla_a \nabla_b \rho_{mv} + \dots \right]$$

↓ Even-Even + integrate over  $\int d\vec{r}$

$$\rho_x(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \approx \Pi_0(k_F r) \rho_x(\vec{R}) + \frac{k_F^2}{6} \Pi_2(k_F r) \left[ \frac{1}{4} \nabla^2 \rho_x(\vec{R}) - \zeta_x(\vec{R}) + \frac{3}{5} k_F^2 \rho_x(\vec{R}) \right]$$

$$k_F = k_F(\vec{R}) = \left[ \frac{3 \Pi^2}{2} \rho(\vec{R}) \right]^{1/3}$$

$$\vec{S}_x(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \approx -\frac{i}{2} \Pi_1(k_F r) \vec{r} \times \vec{J}_x(\vec{R})$$

PSA-DME (w/  $g(\vec{r}, k) = \theta(k_F - k)$ )

$$\Pi_0 = \Pi_1 = \Pi_2 = \frac{3 j_1(k_F r)}{k_F r} = \rho_{SL}(k_F r)$$

Negele-Vautherin DME

$$\Pi_0 = \rho_{SL}(k_F r) \quad \Pi_2 = 1.05 \frac{j_3(k_F r)}{(k_F r)^3}$$

$$\Pi_1 = j_0(k_F r)$$



# HF energy for realistic (e.g., chiral EFT) interaction

$$E_{\text{int}}^{\text{HF}} = \frac{1}{2} \sum_{\substack{\sigma_1, \sigma_2 \\ \tau_1, \dots, \tau_4}} \prod_{i=1}^4 \int d\vec{r}_i \langle \vec{r}_1 \sigma_1 \tau_1, \vec{r}_2 \sigma_2 \tau_2 | \hat{V}_{A_{12}} | \vec{r}_3 \sigma_3 \tau_3, \vec{r}_4 \sigma_4 \tau_4 \rangle \rho(\vec{r}_3 \sigma_3 \tau_3, \vec{r}_1 \sigma_1 \tau_1) \rho(\vec{r}_4 \sigma_4 \tau_4, \vec{r}_2 \sigma_2 \tau_2)$$

$$= \frac{1}{2} \text{Tr}_{\sigma_2}^{(1)} \text{Tr}_{\sigma_2}^{(2)} \int \prod_{i=1}^4 d\vec{r}_i \langle \vec{r}_1 \vec{r}_2 | \hat{\mathcal{V}}^{(12)} | \vec{r}_3 \vec{r}_4 \rangle \rho_{\sim}^{(1)}(\vec{r}_3, \vec{r}_1) \rho_{\sim}^{(2)}(\vec{r}_4, \vec{r}_2) \quad (\hat{\mathcal{V}} = \hat{V}_{A_{12}})$$

$\nearrow$  Matrix on spin/isospin space  $1 \otimes 2$      
  $\nearrow$  Matrix in spin/isospin space of particle 1     
  $\nearrow$  matrix in spin/isospin space of particle 2

\* hereafter drop  $1 \otimes 2$  + particle 1 + 2 labels unless absolutely needed

$$\text{Now, } \langle \vec{r}_1 \vec{r}_2 | \hat{\mathcal{V}} | \vec{r}_3 \vec{r}_4 \rangle = \delta(\vec{R} - \vec{R}') \langle \vec{r} | \hat{\mathcal{V}} | \vec{r}' \rangle$$

$$\Rightarrow E_{\text{int}}^{\text{HF}} = \frac{1}{2} \text{Tr}_{\sigma_2}^{(1)} \text{Tr}_{\sigma_2}^{(2)} \int d\vec{R} d\vec{r} d\vec{r}' \langle \vec{r} | \hat{\mathcal{V}} | \vec{r}' \rangle \rho_{\sim}(\vec{R} + \frac{\vec{r}}{2}, \vec{R} + \frac{\vec{r}'}{2}) \rho_{\sim}(\vec{R} - \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}'}{2})$$

\* Where A.S. potential in coordinate space given as

$$\langle \vec{r} | \hat{\mathcal{V}} | \vec{r}' \rangle = \int \frac{d\vec{p}}{(2\pi)^3} \frac{d\vec{p}'}{(2\pi)^3} e^{i\vec{p} \cdot \vec{r}'} e^{-i\vec{p}' \cdot \vec{r}} \langle \vec{p}' | \hat{V}(1 - P_0 P_z P_0) | \vec{p} \rangle$$

$$= \int \frac{d\vec{p} d\vec{p}'}{(2\pi)^6} e^{i\vec{p} \cdot \vec{r}' - i\vec{p}' \cdot \vec{r}} (\langle \vec{p}' | \hat{V} | \vec{p} \rangle - \langle \vec{p}' | \hat{V} P_0 P_z | \vec{p} \rangle)$$

Note: Why bother w/  $\vec{p}$ -space? Because modern chiral EFT potentials most naturally given in momentum space.

General form of  $\hat{V}$  as in, say, chiral EFT

$$\begin{aligned} \langle \vec{p}' | \hat{V} | \vec{p} \rangle &= [V_c + \tau_1 \tau_2 W_c] + [V_s + \tau_1 \tau_2 W_s] \vec{\sigma}_1 \cdot \vec{\sigma}_2 + [V_T + \dots] \vec{\sigma}_1 \cdot \vec{q} \sigma_2 \cdot \vec{q} \\ &+ [\tilde{V}_c + \tau_1 \tau_2 \tilde{W}_c] + [\tilde{V}_s + \dots] \vec{\sigma}_1 \cdot \vec{\sigma}_2 + [\tilde{V}_T + \dots] \vec{\sigma}_1 \cdot \vec{k} \sigma_2 \cdot \vec{k} \\ &+ [V_{LS} + \tau_1 \tau_2 W_{LS}] \frac{i}{2} (\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{q} \times \vec{k} \end{aligned}$$

where  $\vec{q} = \vec{p}' - \vec{p}$  and  $\{V_c, W_c, \dots\}$  depend on  $\vec{q}$  only  
 $\vec{k} = \frac{\vec{p} + \vec{p}'}{2}$   $\{\tilde{V}_c, \tilde{W}_c, \dots\}$  depend on  $\vec{k}$  only

If this is obscure to you, think of a term in the Minnesota potential without a  $P_r$  factor

e.g.  $V(r) \quad \langle \vec{p}' | V | \vec{p} \rangle = \int d^3r V(r) e^{i\vec{r} \cdot (\vec{p}' - \vec{p})}$  only depends on  $\vec{q}$

\* Now consider a term  $\underline{V(r)P_r}$ :  $\langle \vec{p}' | V P_r | \vec{p} \rangle = \int d^3r \langle \vec{p}' | \vec{r} \rangle V(r) \langle \vec{r} | P_r | \vec{p} \rangle = \int d^3r \langle \vec{p}' | \vec{r} \rangle V(r) \langle -\vec{r} | \vec{p} \rangle$   
 $= \int d^3r e^{-i\vec{r} \cdot (\vec{p}' + \vec{p})} V(r)$  only depends on  $\vec{k} = \frac{\vec{p} + \vec{p}'}{2}$

\* Likewise, you can expand

$$\begin{aligned} \langle \vec{p}' | \hat{V} P_r P_r | -\vec{p} \rangle &= [V_c^x + \dots] + \dots + [V_T^x + \tau_1 \tau_2 W_T^x] \vec{\sigma}_1 \cdot \vec{k} \sigma_2 \cdot \vec{k} \\ &+ [\tilde{V}_c^x + \dots] + \dots + [\tilde{V}_T^x + \dots] \vec{\sigma}_1 \cdot \vec{q} \sigma_2 \cdot \vec{q} \end{aligned}$$

where  $\{V_c^x, \dots, V_T^x\}$  depend only on  $\vec{k} = \frac{\vec{p}' + \vec{p}}{2}$

$\{\tilde{V}_c^x, \dots, \tilde{V}_T^x\}$  depend only on  $\vec{q} = \vec{p}' - \vec{p}$

Note: The  $\{V_i^X, W_i^X, \tilde{V}_i^X, \tilde{W}_i^X\}$  can be expressed as linear combos of the  $\{V_i, W_i, \tilde{V}_i, \tilde{W}_i\}$

claim: Making use of the above components of NN-interaction, plus the following

$$P_0(\vec{r}_1, \vec{r}_2) \equiv \text{Tr}_{\sigma z} \tilde{P}_0(\vec{r}_1, \vec{r}_2)$$

$$P_1(\vec{r}_1, \vec{r}_2) \equiv \text{Tr}_{\sigma z} [P_2(\vec{r}_1, \vec{r}_2) \tilde{\tau}_z]$$

$$\vec{S}_0(\vec{r}_1, \vec{r}_2) = \text{Tr}_{\sigma z} [P_2(\vec{r}_1, \vec{r}_2) \vec{\sigma}]$$

$$\vec{S}_1(\vec{r}_1, \vec{r}_2) = \text{Tr}_{\sigma z} [P_2(\vec{r}_1, \vec{r}_2) \vec{\sigma} \tilde{\tau}_z]$$

For even-even  
Nuclei for  
Simplicity



$$E_H = \frac{1}{2} \sum_{\alpha=0,1} \int d\vec{R} d\vec{r} \left[ P_\alpha(\vec{R} + \frac{\vec{r}}{2}) P_\alpha(\vec{R} - \frac{\vec{r}}{2}) \Gamma_c^{\alpha t}(\vec{r}) + \vec{r} \cdot \vec{S}_\alpha(\vec{R} + \frac{\vec{r}}{2}) P_\alpha(\vec{R} - \frac{\vec{r}}{2}) \Gamma_{L_S}^{\alpha t}(\vec{r}) \right]$$

$$E_F = -\frac{1}{2} \sum_{\alpha=0,1} \int d\vec{R} d\vec{r} \left[ P_\alpha^2(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \Gamma_c^{\alpha t}(\vec{r}) - \vec{S}_\alpha^2(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \Gamma_S^{\alpha t} \right. \\ \left. + S_\alpha^\alpha(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) S_\alpha^\beta(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \nabla_\alpha \nabla_\beta \Gamma_T^{\alpha t}(\vec{r}) \right. \\ \left. + \lambda \Gamma_{L_S}^{\alpha t}(\vec{r}) \vec{S}_\alpha(\vec{R} + \frac{\vec{r}}{2}, \vec{R} - \frac{\vec{r}}{2}) \cdot (\vec{r} \times \vec{\nabla}_{i2}) P_\alpha(\vec{r}_1, \vec{r}_2) \right]$$

where:

$$\Gamma_\alpha^k(\vec{r}) \equiv V_\alpha^k(\vec{r}) - \tilde{V}_\alpha^k(\vec{r}) \quad t=0 \quad \Gamma_\alpha^{\alpha t}(\vec{r}) \equiv V_\alpha^{\alpha X}(\vec{r}) - \tilde{V}_\alpha^{\alpha X}(\vec{r}) \quad t=0 \\ \equiv W_\alpha^k(\vec{r}) - \tilde{W}_\alpha^k(\vec{r}) \quad t=1 \quad \equiv W_\alpha^{\alpha X}(\vec{r}) - \tilde{W}_\alpha^{\alpha X}(\vec{r}) \quad t=1$$



$$C_x^g(m) = \sum_{n=0}^2 C_{x,n}^g(m) \quad g \in \{PP, PT, PDP, \dots\}$$

$$C_{x,n}^g(m) = \underbrace{\alpha_0^g(n,t,m)} + \sum_{j=1}^2 \underbrace{\alpha_j^g(n,t,m)}_j \underbrace{F_j(n,m)}_j$$

rational Polynomial in  $m$ 
Non-analytic in  $m$  due to finite range

e.g. LO ( $1\pi$ -exchange):

$$F_1(0,m) = \log(1+4m^2)$$

$$F_2(0,m) = \text{Arctan } 2m$$

NLO ( $2\pi$ -exchange):

$$F_1(1,m) = \left[ \log(1+2m^2+2m\sqrt{1+m^2}) \right]^2$$

$$F_2(1,m) = \sqrt{1+m^2} \log(1+2m^2+2m\sqrt{1+m^2})$$